# 'Reionization' Experiment: PAPER

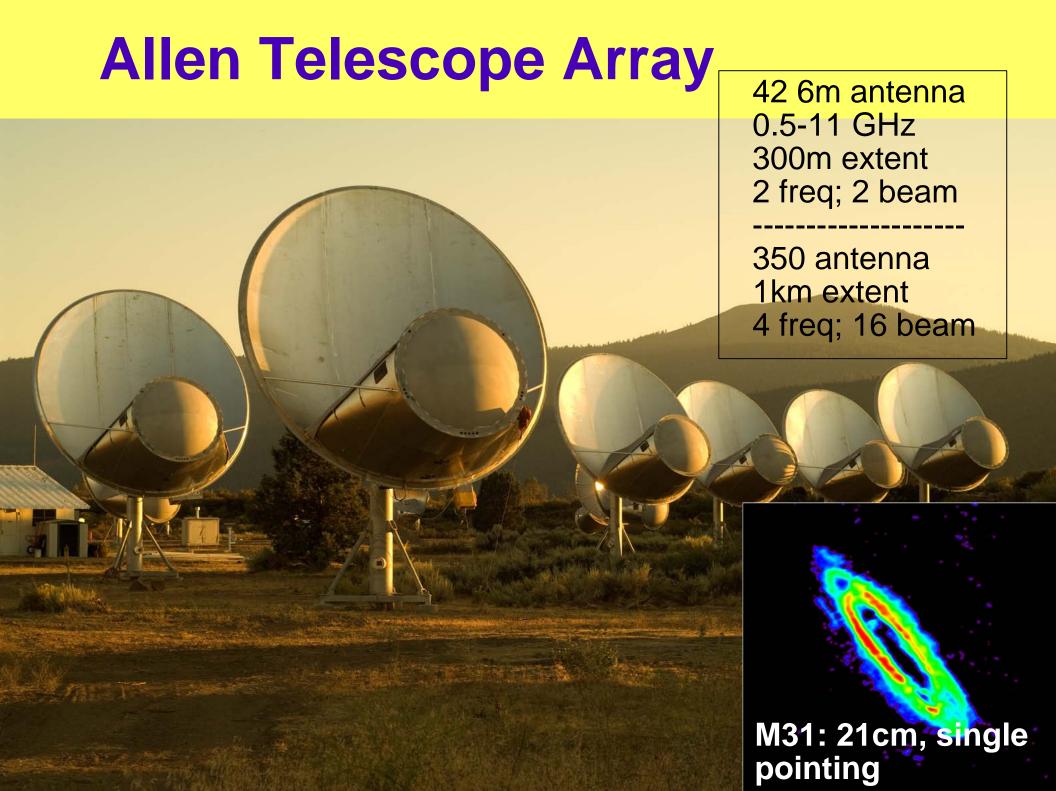
#### D. C. Backer

# Astronomy Department University of California, Berkeley

Berkeley Radio Astronomy Laboratory programs introduction "innovation & exploration"

PAPER: Precision Array to Probe the Epoch of Reionization

- •Epoch of Reionization (also next talk by Xuelei Chen)
- •Green Bank test array: PGB
- •Western Australia deployment: PWA
- Future



# CARMA: Combined Array for Research in Millimeter Astronomy



# CASPER



CENTER FOR ASTRONOMY SIGNAL PROCESSING AND ELECTRONICS RESEARCH

The goal of CASPER is to streamline and reduce the current radio astronomy instrumentation design flow through the development of an open-source, platform-independent design approach. This incorporates reconfigurable, modular, easily upgradable hardware with standard, parameterized design libraries that abstract away the underlying details of the system. Design simplifies to creation of block diagrams of components from standard libraries and the designer.

**HARDWARE:** ADC, FPGA Board, Switch, CPU

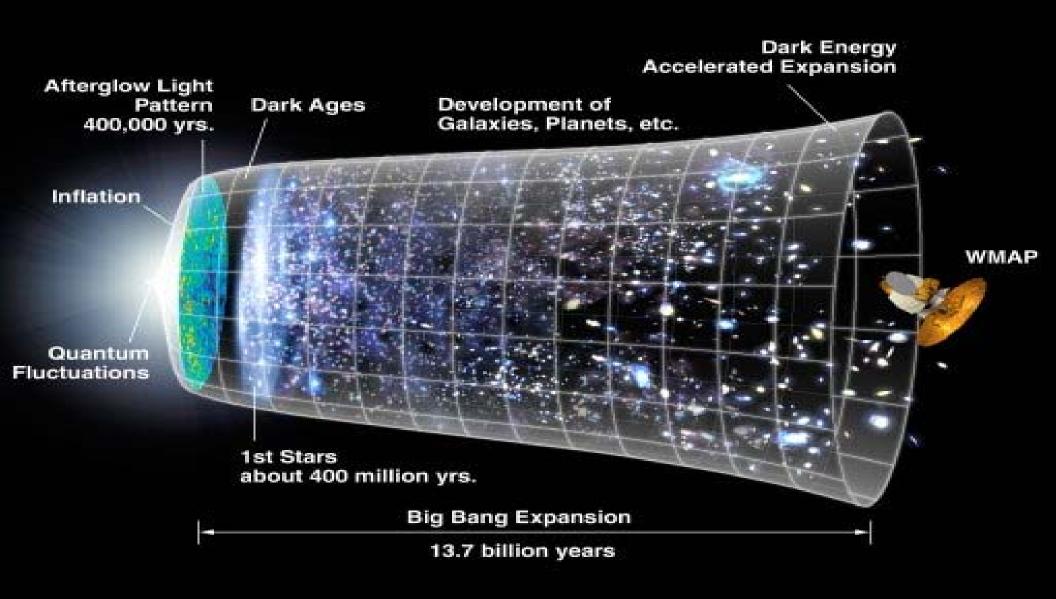
<u>DESIGNS:</u> Spectrometer, Correlator, Beamformer, Pulsarometer, Realtime Imager

# PRECISION ARRAY TO PROBE THE EPOCH OF REIONIZATION

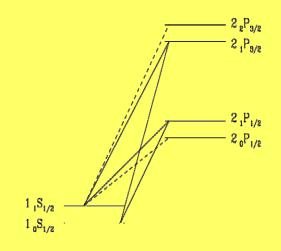
#### PAPER Team:

R. Bradley (Co PI), E. Mastrantonio, C. Parashare, N. Gugliucci, D. Boyd, P. Reis (*NRAO & UVA*); A. Parsons, M. Wright, D. Werthimer, CASPER group (*UC Berkeley*); D. Herne, M. Lynch (*Curtin Univ*); C. Carilli, A. Datta (*NRAO Socorro*); J. Aguirre (*U Colorado*)

Our experiment is working toward detection of this excited 21cm line in the "swiss cheese" like 21cm brightness temperature structure: image the sky at many frequencies, difference in angle and in frequency (red shift or time), form a statistical summary to find signal.



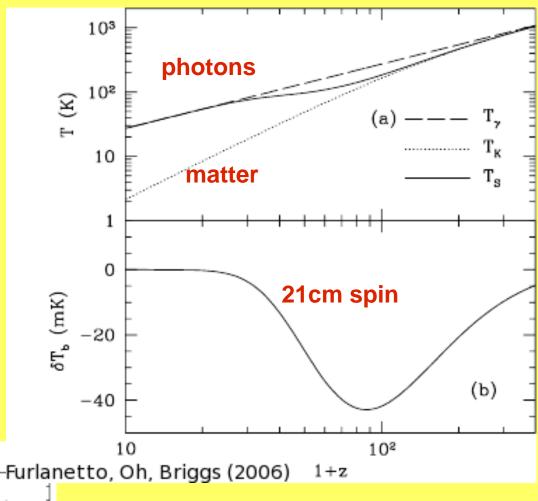
## 21 cm Spin Temp

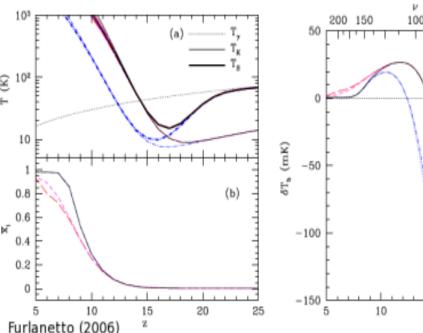


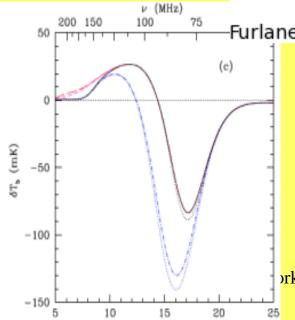
Wouthuysen Field Effect

$$T_{\gamma} \propto (1+z)$$

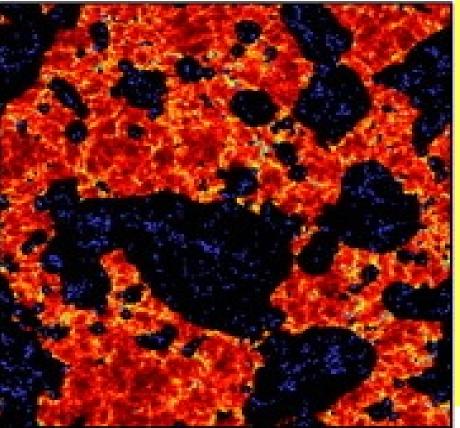
$$T_K \propto (1+z)^2$$





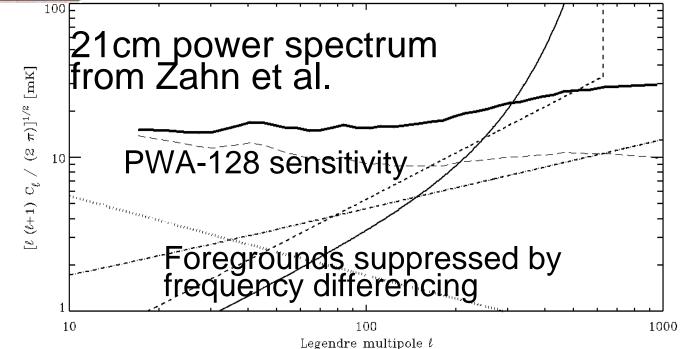


orkshop



# Hydrogen Reionization Simulation and Power Spectrum

Morphology of 21cm bubbles in late stage, z~7, of reionization when the neutral fraction is approximately 50% (Zahn et al. 2007ApJ...654...12Z). The size of the box is 66 Mpc h<sup>-1</sup> on a side and 0.25 Mpc h<sup>-1</sup> deep.



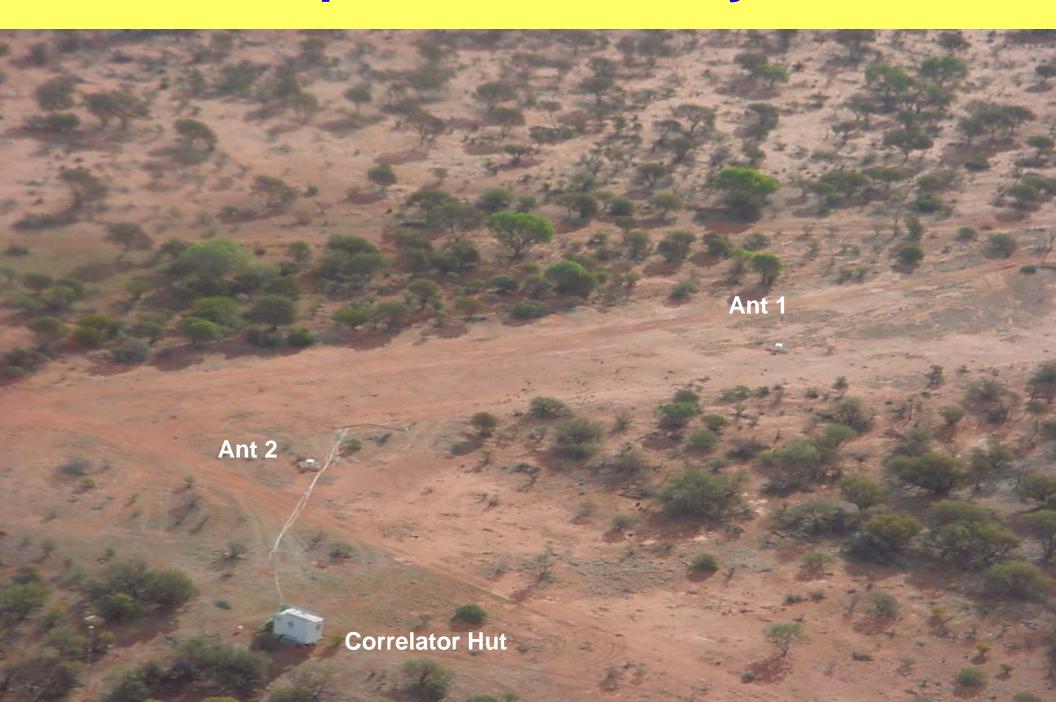
#### "PAPER" PROGRESS

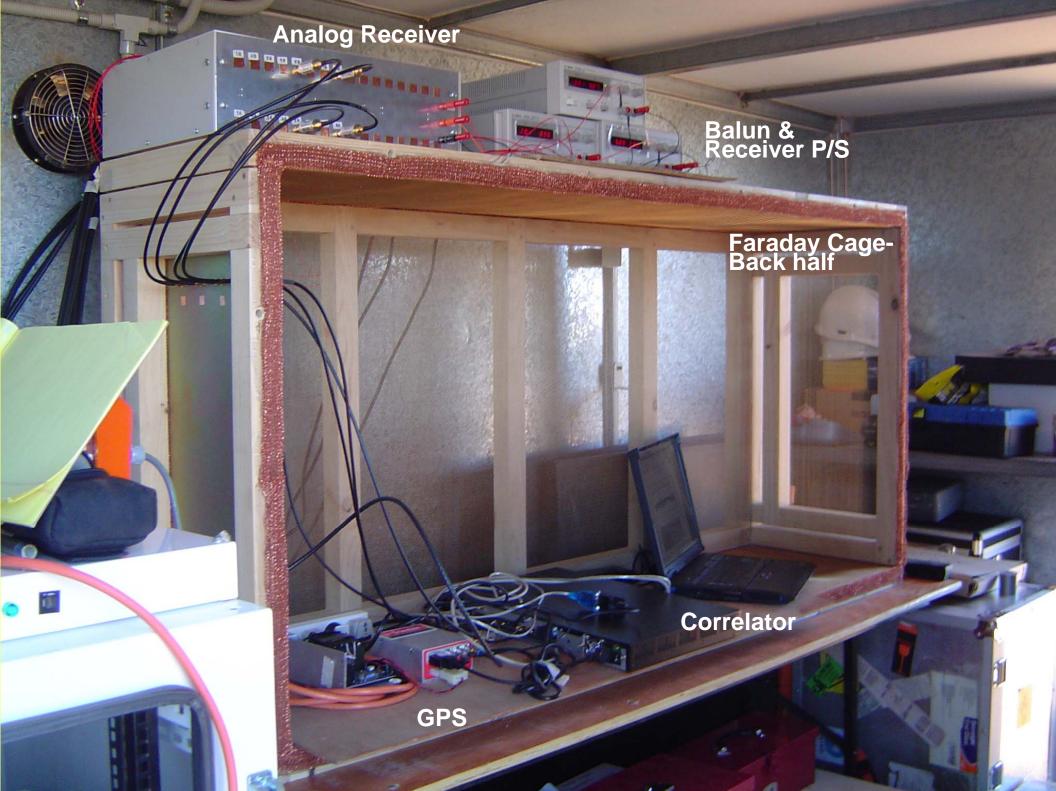
- Start in 2004; NSF funding: 2005-2007 correlator development grant; 2006-2008 experiment "starter" grant, including WA deployment; other funding via parallel projects (correlator, FASR) and Carilli MPG award; pending 2008-2011 NSF.
- PAPER in Green Bank: PGB. This has evolved from 2-antenna interferometer in 2004 August to 8-antenna array in 2006; 16-antenna array 2008 May; also, single-antenna test facility.
- PAPER in Western Australia: PWA. 4-dipole array deployed: 2007 July.
- PGB 8-antenna 2008 March with revised design

#### John Richards-lease holder; Ron Beresford, CSIRO; DB



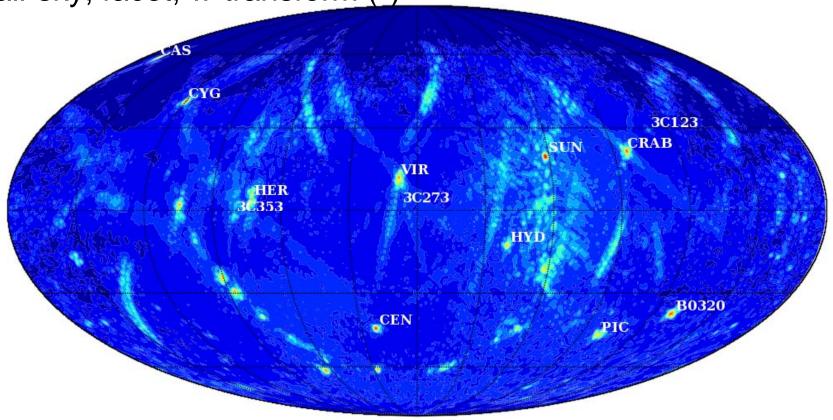
## PWA-4 —Top Shed, Boolardy Station





## PAPER in Western Australia—2007 July

125-190 MHz 4 Dipole 24 hour integration MFS, all-sky, facet, w-transform (\*)

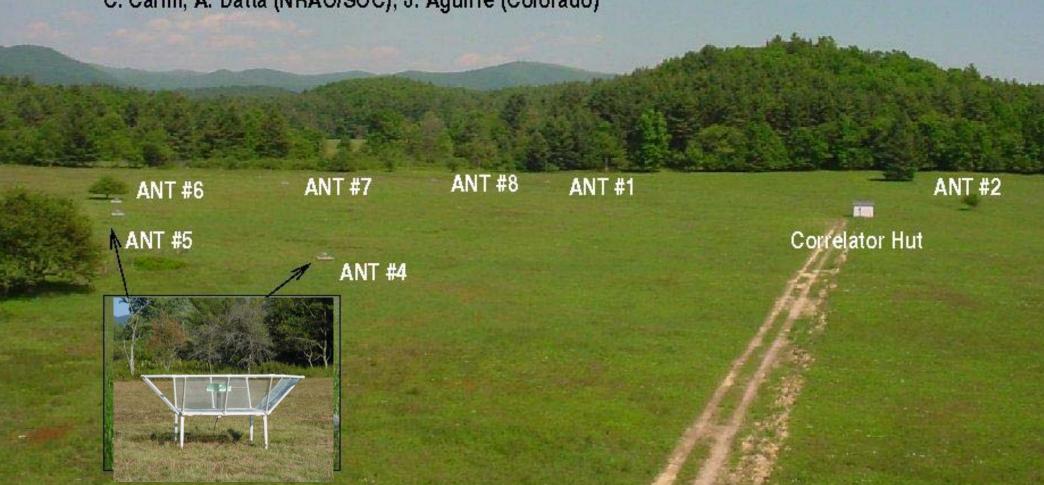


RMS away from strong sources: ~1 Jy/~1 K (\*) AIPY – Astronomical Imaging in Python – Aaron Parsons

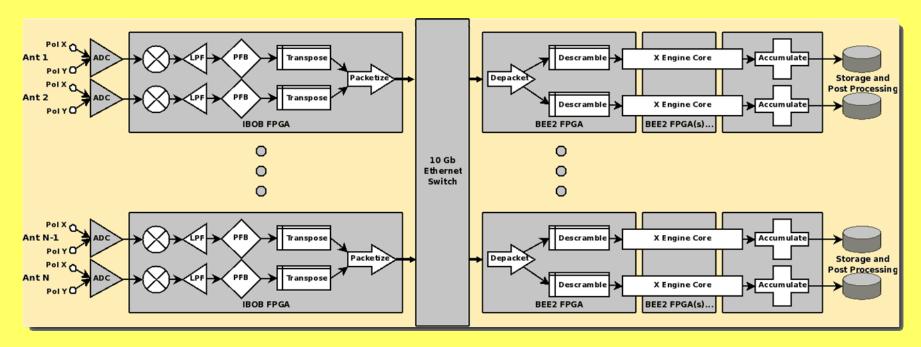
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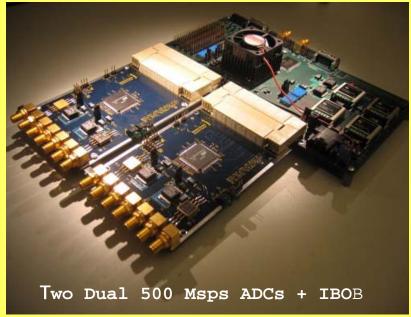
#### GALFORD MEADOW -- NRAO: GREEN BANK, WV

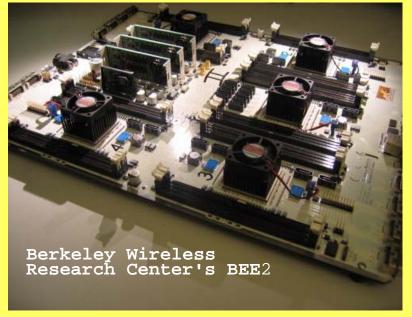
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#### PAPER/CASPER Packetized Correlator



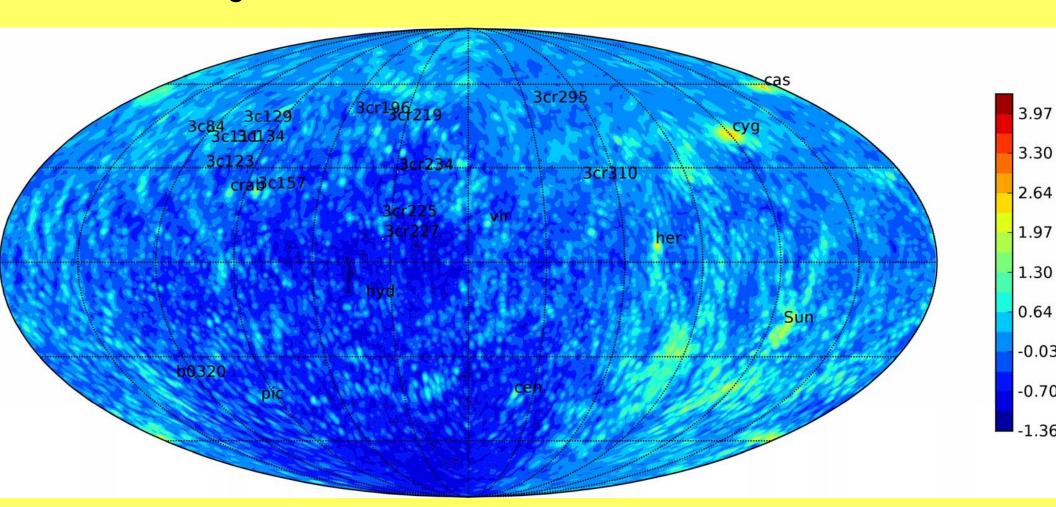




2008 Apr 21 Sino - US Workshop 15

## PAPER in Green Bank—2008 Mar

130-170 MHz7 Dipole24-hour integration



## **SUMMARY**

- Step by step approach successful
- Green Bank test array essential
- AIPY and related calibration/imaging just starting: beam fitting; polarization soon
- Gearing up for PWA-32 deployment 2008 Sep
- Funding looks good for buildout to PWA-128 in years ahead: power spectrum detectability dependent on configuration, foreground removal, other systematics.
- Long term: 100M USD effort decision middecade.

